



Energetic electron precipitations associated with ULF Pc5 modulations during substorms using LF/VLF standard radio waves

Hiroyo Ohya*⁽¹⁾, Takuya Miyashita⁽¹⁾, Fuminori Tsuchiya⁽²⁾, Kazuo Shiokawa⁽³⁾, Yoshizumi Miyoshi⁽³⁾, and Martin Connors⁽⁴⁾

(1) Chiba University, Chiba, 263-8522, e-mail: ohya@faculty.chiba-u.jp; aeka2354@chiba-u.jp

(2) Tohoku University, Sendai, Japan; e-mail: tsuchiya@pparc.gp.tohoku.ac.jp

(3) Nagoya University, Nagoya, Japan; e-mail: shiokawa@nagoya-u.jp; miyoshi@isee.nagoya-u.ac.jp

(4) Athabasca University, Athabasca, Canada; e-mail: martinc@athabascau.ca

Periodic electron precipitations into the lower ionosphere are often observed by ground-based riometers as cosmic noise absorption [1]. The period of the variations is from a few seconds to several minutes associated with ultra low frequency (ULF) magnetic pulsations [2]. The possible mechanism is that the ULF waves in the magnetosphere modulate the precipitations of energetic electrons [3], which means the electron precipitations and the ULF waves have same periods. In the ULF waves, Pc5 (150-600 s) pulsations and electron precipitations are simultaneously observed in the morning sector (6-12 MLT) at the auroral latitudes [4]. In this study, we investigate the precipitations of energetic electrons into the lower ionosphere during substorms of 27 April, 2017 and 7-8 September, 2017, using a network of very low frequency (VLF)/low frequency (LF) subionospheric transmitter signals.

As for the substorm of 27 April, 2017, the transmitters are NAA (USA, 24.0 kHz, $L = 2.88$), NLK (USA, 24.8 kHz, $L = 2.88$), NDK (USA, 25.2 kHz, $L = 2.98$) and WWVB (USA, 60.0 kHz, $L = 2.26$), while the receivers are ATH (Athabasca, $L = 4.31$) and PKR (Poker Flat, $L = 5.95$). As for the substorm of 7-8 September, 2017, the transmitters are NRK (Iceland, 37.5 kHz, $L = 4.05$) and NPM (Hawaii, 21.4kHz, $L = 1.15$), while the receivers are NYA (Ny-Ålesund, $L=17.2$) and PKR. The substorm occurred at 05:35 UT (around midnight in American sector), 27 April, 2017. Based on the wavelet analysis, we found that the LF/VLF intensities for the NDK-ATH, WWVB-ATH, and NLK-ATH paths show oscillations with a period of 4-6 minutes during the substorm growth phase (05:00-06:00 UT). Magnetic field variations at ATH also show the same period, indicating that the radio intensity variations with the period of 4-6 minutes would be caused by ULF Pc5 modulations.

We also analyzed the substorm occurred at 19:40 UT on 7 September, 2017. Variations in the radio intensities (3-4 dB) for NRK-NYA, NRK-PKR and NPM-NYA paths due to precipitations were seen during 23:20 UT, 7 September - 01:00 UT, 8 September, 2017. Based on the wavelet analysis, the radio intensities for above 3 paths show oscillations with periods of 30-40 minutes. Magnetic field variations along the subionospheric propagation paths, the IMF Bz, and AE index also showed the same periods of 30-40 minutes, which could be caused by the occurrence of the substorm injection itself. In the presentation, we will discuss the cause of these radio variations in detail.

1. T. Motoba, K. Takahashi, J. Gjerloev, S. Ohtani, and D. K. Milling, "The role of compressional Pc5 pulsations in modulating precipitation of energetic electrons", *J. Geophys. Res. Space Physics*, **118**, 2013, pp. 7728–7739, doi:10.1002/2013JA018912.

2. J. R. Barcus, and T. J. Rosengerg, "Observations on the spatial structure of pulsating electron precipitation accompanying low frequency hydromagnetic disturbances in the auroral zone", *J. Geophys. Res.*, **70**, 1965, pp. 1707–1716.

3. F. V. Coroniti, and C. F. Kennel, "Electron precipitation pulsations", *J. Geophys. Res.*, **75**, 1970, pp.1279–1289.

4. J. B. Baker, E. F. Donovan, and B. J. Jackel, "A comprehensive survey of auroral latitude Pc5 pulsation characteristics", *J. Geophys. Res.*, **108**(A10), 1384, 2003, doi:10.1029/2002JA009801.